



Optical Aperture Synthesis

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Outline

- Introduction
 - Objectives and Benefits
 - Background
 - Principles
- Programme
 - Risk Analysis
 - Risk Mitigation

Objectives and Benefits

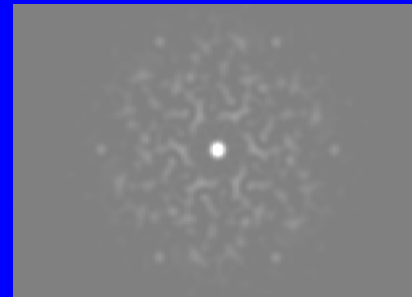
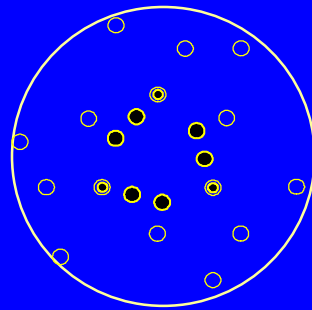
- Array of small apertures
 - Lightweight
 - Deployable (terrestrial)
 - High-resolution
 - Long range recognition and ID
 - >10x range or 10x resolution
- Distributed, self-calibrating system
 - Good-quality images

Objectives and Benefits

- Synthetic imaging can give high-resolution images through small apertures

Angular
Resolution

$$\alpha = 1.22 \frac{\lambda}{D}$$



For OAS
 D is max
baseline
- not
aperture
diameter

Background

- Solicitation for Research Proposal from EMRS DTC dated 21 May 2003
 - Risk analysis
 - Literature survey
 - Mitigation strategy

Principles

- Principles well known
 - Used in astronomy at optical-, radio-, mm- λ
- Unused in military
 - High reward, high-risk
- Why high risk?
 - Why different from astronomy?
 - Lower data volume
 - Snapshot images needed
 - Extended, cluttered scenes
 - Less benign environment

New features in Surveillance

- Crowded space on platform
 - Bulk optics beam transport not feasible
→ use waveguides
- Vibrating, manoeuvring platform
 - Path-length compensation
- Active illumination
 - Increase flux level
- Thermo/mechanical load
 - Use AO correction

Risk Analysis

- 3 **Critical** Risks
 - Must be resolved to achieve anything useful. Fundamental.
- 6 **Severe** Risks
 - Must be resolved to achieve useful results but incremental nature.
- 5 **Important** Risks
 - Must be resolved to obtain good performance.

Risks

Critical

- Pointing and beam-feeding
- OPD equalisation
- Instrument calibration

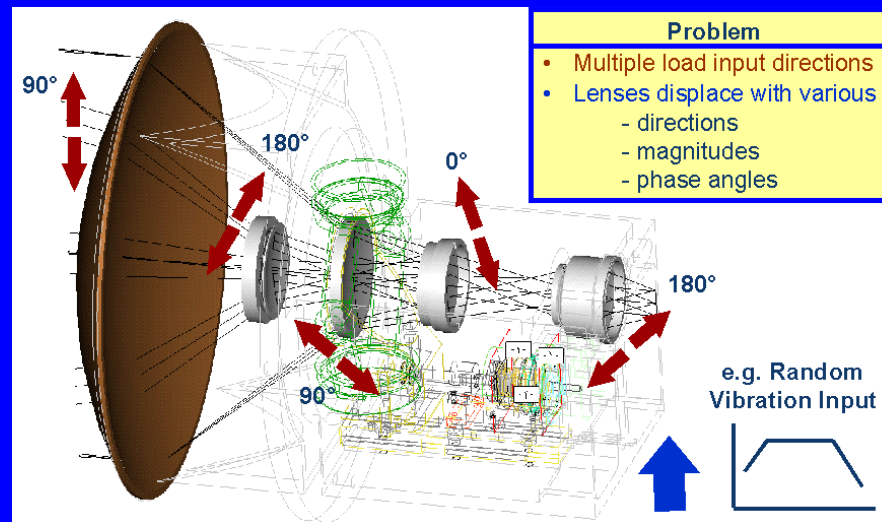
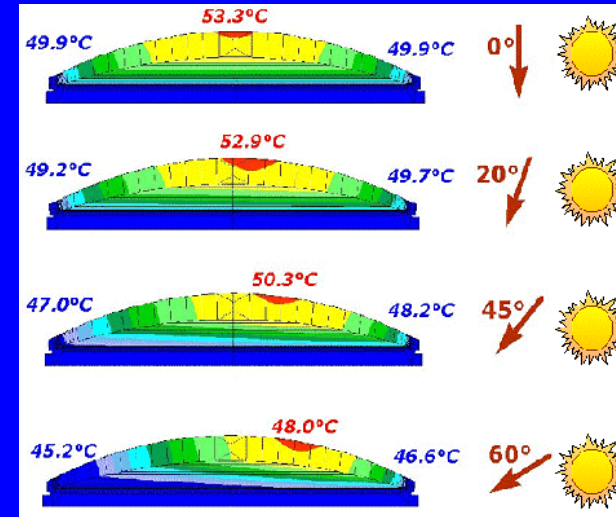
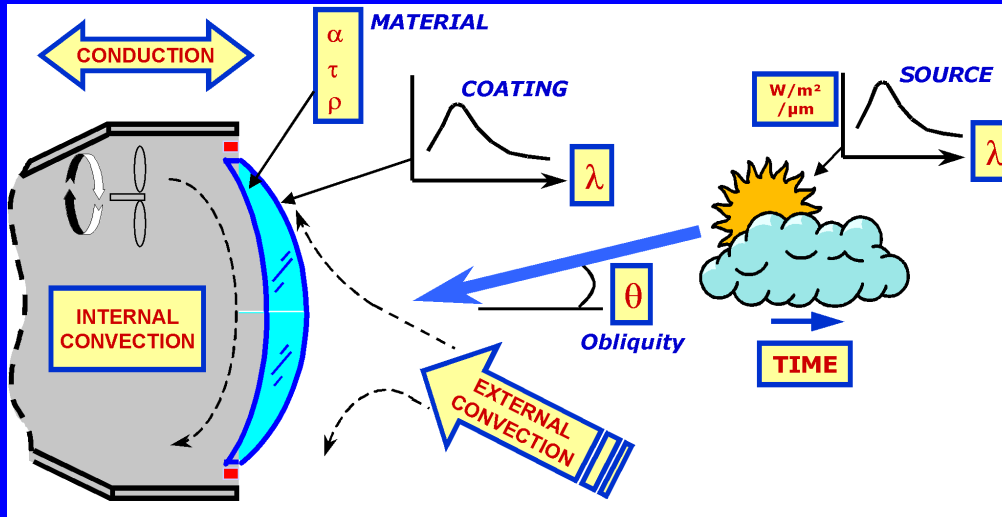
To get interference:

- Must overlap images from each collector accurately
- Must operate within coherence length
- Must achieve real-time calibration

Critical Risk Treatment

- Pointing & feeding
 - SNR implications
 - Platform constraints
- OPD control
 - Internal control OK
 - External calibration
- Instrument calibration
 - RSC, Fourier telescope

Thermo/mechanical load



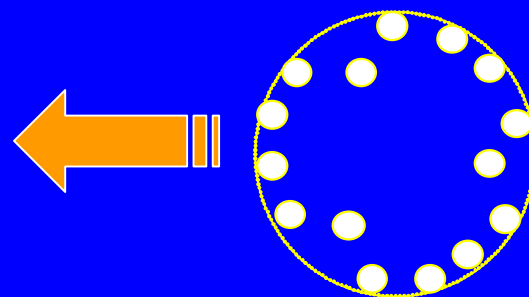
Illustrations courtesy Thales Optics UK

Attention to design

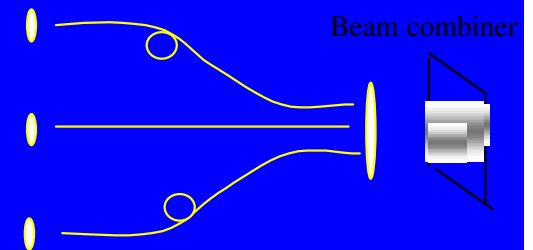
Wish	1975	2004
Greater Boresight Stability Higher Tolerance of Vibration Capable of handling Higher Loads Smaller space envelope Lower Mass	0.25 to 0.50mradians Low tolerance 25 to 100g Possible within reason Possible	20 to 50 μ radians 10g rms random 50 to 150g Small & complex Reduce by 3-5% per annum
Result	Make & see if it works	Modeling before manufacture



May 2004



Collector array



Waveguide beam transport

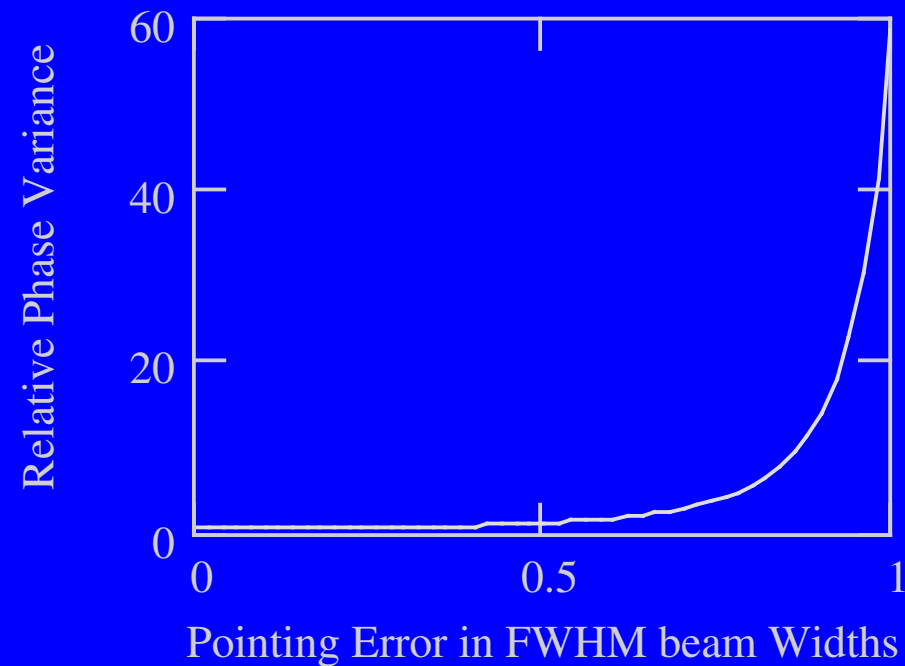
Pointing & Feeding

- Vibration within optical subsystem
 - μ -radian errors
- Flexing airframe
 - milli-rad errors
 - inertial gimbals - 50Hz
 - adaptive control
 - Offset to flexure errors with long time constant
- ‘Static’ boresight errors

Pointing & Feeding

- Images must overlap to get interference
- Fibre feed must be very accurate
- Phase accuracy

$$\sigma_{\phi}^2 = \frac{1}{2n|J(\xi)|^2}$$

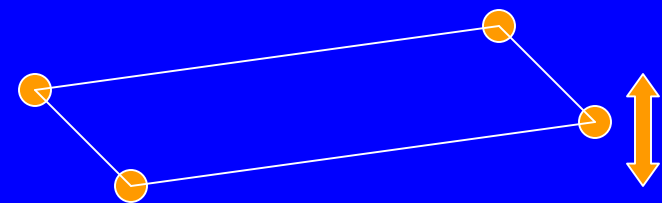


OPD correction

- Within coherence length
 - Required for passive operation
 - Needed for interference
- Active illumination
 - Narrow band
 - Relaxed OPD control
 - Speckle effects
- Within $\lambda/4$ (active or passive)
 - Needed for high image quality

Redundant Spacings Calibration

- Repeated measurement of a Fourier component should give same result
 - Differences relate to instrument/atmosphere and not to target
- Gives model-independent calibration of OAS array
 - Direct matrix inversion or use as adaptive instrument



Fourier Telescopy

- Fourier integral underlies imaging

$$\int dr o(r) e^{-ikr.\xi} \Rightarrow$$

$$\int dr o(r) (\cos(kr.\xi) - i \sin(kr.\xi))$$

- Coherent projection of fringes
- Modulation to get Sin & Cos
- Integration

- Mathematical image reconstruction
- Large area detector for sensitivity
- RSC for calibration and anisoplanatism

Active imaging

- Speckled images
 - Unfamiliar appearance
 - Speckle mitigation
- Betrays observation
 - Increased range
 - Increased opportunity
- Speckle statistics?
 - Geometric data?
- ‘Holographic opportunity’
 - Phase-sensitive detection?

Other Possibilities

- Parametric imaging
 - May not require full imaging facility
- Speckle statistics
 - Shape of target
 - Surface finish
- Target statistics
 - Number and relationship of facets
 - Number and relationship of retro-reflections

Optical Aperture Synthesis

- High risk but big payoff
 - Parametric imaging
 - Technically challenging
 - Requires staged approach
- Greatly-increased range/resolution
 - Recognition and ID based on key features?
 - Challenges not intractable
 - Manage risk and cost/benefit